## SMS-618, Particle Dynamics, Fall 2003 (E. Boss, last updated: 10/22/2003) Single Particle Dynamics

## Solute transport to/from a sinking cell

The conservation equation for a solute is given by (as in lecture 4):

$$
\begin{equation*}
\frac{\partial C}{\partial t}+\bar{u} \nabla C=\nabla(D \nabla C), \nabla \equiv \partial / \partial x+\partial / \partial y+\partial / \partial z \tag{1}
\end{equation*}
$$

where $\vec{u}$ is the velocity field ( $\left[\mathrm{L} \mathrm{T}^{-1}\right]$ )., and $D$ the diffusion coefficient $\left(\left[\mathrm{L}^{2} \mathrm{~T}^{-1}\right]\right)$. We wish to solve this equation for a particle. Let's assume a spherical particles (radius, $r[\mathrm{~L}]$ ) sinking steadily in the fluid $\left(\partial C / \partial t=0, \vec{u}_{\text {particle }}=w_{s}\right)$ and a constant diffusion coefficient $D=$ const. around the cell:

$$
\begin{align*}
& \vec{u} \cdot \nabla C=D \nabla^{2} C, \quad \nabla C=\hat{r} \frac{\partial C}{\partial r}+\frac{\hat{\theta}}{r} \frac{\partial C}{\partial \theta}+\frac{1}{r \sin \theta} \frac{\partial C}{\partial \phi} \\
& \nabla^{2} \equiv \frac{1}{r^{2}} \frac{\partial}{\partial r}\left(r^{2} \frac{\partial}{\partial r}\right)+\frac{1}{r^{2} \sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta \frac{\partial}{\partial \theta}\right)+\frac{1}{r^{2} \sin ^{2} \theta} \frac{\partial}{\partial \phi^{2}} \tag{2}
\end{align*}
$$

By invoking azimuthal symmetry of both velocity and concentration fields, $\partial / \partial \phi=0$.
Equation (2) is non-dimensionalized as follows: $\mathrm{C}^{*}=\mathrm{C} /\left|\mathrm{C}_{\infty}-\mathrm{C}_{0}\right|, \mathrm{u}^{*}=\mathrm{u} / \mathrm{w}_{\mathrm{s}}, \mathrm{r}^{*}=\mathrm{r} / \mathrm{r}_{0}, 0$ denoting the particles perimeter. The equation for the non-dimensional variables is:

$$
\begin{equation*}
P e\left(\vec{u}^{*} \cdot \nabla C^{*}\right)=\nabla^{2} C^{*}, \quad P e \equiv \frac{w_{s} r_{0}}{D} . \tag{3}
\end{equation*}
$$

The non-dimensional Peclet number, $P e=R e^{*} D / v$, is a measure of the relative domination of advection over diffusion in terms of solute flux to/away of the sinking sphere.

The ratio of the flux to/from the particle in the presence of motion to the flux when there is no motion is another non-dimensional parameter, the Sherwood number, Sh :

$$
\begin{equation*}
S h=\frac{-D \int_{A} \hat{n} \cdot \nabla C d A}{4 \pi r_{0} D\left(C_{\infty}-C_{0}\right)} \tag{4}
\end{equation*}
$$

(the diffusive solution was derived in lecture 4).
A relationship between $S h$ and $P e$ for $0.001<\mathrm{Pe}<5000$ that is consistent with analytical solutions for small and large $P e$ is (Clift et al., 1978, Karp-Boss et al., 1996):

$$
\begin{equation*}
S h=\frac{1}{2}\left(1+(1+2 P e)^{\frac{1}{3}}\right) \tag{5}
\end{equation*}
$$

## Solute transport to/from a cell in a turbulent fluid:

In the case of a turbulent fluid, the flow around small particles can be characterized based on the energy dissipation rate, $\varepsilon$, and the kinematic viscosity, $v$. Assuming that particle to be smaller than the Kolmogorov length scale, $\eta=\left(v^{3} / \varepsilon\right)^{1 / 4}$, the velocity scale around the cell is given by $\mathrm{U}=\mathrm{r}_{0}(\varepsilon / v)^{1 / 2}$. The dependence of $S h$ on $P e \equiv \mathrm{r}_{0}{ }^{2}(\varepsilon / v)^{1 / 2} / \mathrm{D}$ is given by:

$$
S h=\left\{\begin{array}{cc}
1+0.29 P e^{1 / 2} & P e \ll 1  \tag{6}\\
1.014+0.15 P e^{1 / 2}<S h<0.955+0.344 P e^{1 / 3} & 0.01<P e<100 \\
0.55 P e^{1 / 3} & P e \gg 1
\end{array}\right.
$$

## Appendix: Flow around a sinking cell:

The Navier-Stokes equations around a sinking particle have been solved exactly for $\mathrm{Re} \ll 1$. Assuming no flow into the cell, no flow relative to the cell on its perimeter, and U velocity away from the cell (the coordinate system moves with the cell):
$\psi_{\text {Stokes }}=\frac{1}{4}\left(2+\left(\frac{r_{0}}{r}\right)^{3}-3 \frac{r_{0}}{r}\right) U r^{2} \sin ^{2} \theta$,
$u_{\phi}=0, u_{r}=\frac{1}{r^{2} \sin \theta} \frac{\partial \psi}{\partial \theta}=\frac{1}{2}\left(2+\left(\frac{r_{0}}{r}\right)^{3}-3 \frac{r_{0}}{r}\right) U \cos \theta$,
$u_{\theta}=-\frac{1}{r \sin \theta} \frac{\partial \psi}{\partial r}=-\frac{1}{4}\left(4-\left(\frac{r_{0}}{r}\right)^{3}-3 \frac{r_{0}}{r}\right) U \sin \theta$
This solution is correct for $\operatorname{Re}=\mathrm{Ur}_{0} / v=0$. It turns our that as one goes away from the sphere, the neglect of the inertia term is unjustified, and $\mathrm{O}(\mathrm{Re})$ terms have to be incorporated.
The components of the stress on the sphere are:
$\tau_{\phi}=0, \tau_{r}=-p+2 \mu \frac{\partial u_{r}}{\partial r}=-p_{\infty}+\frac{3}{2} \frac{\mu U}{r_{0}} \cos \theta$,
$\tau_{\theta}=\mu r \frac{\partial}{\partial r}\left(\frac{u_{\theta}}{r}\right)+\frac{\mu}{r} \frac{\partial u_{r}}{\partial \theta}=-\frac{3}{2} \frac{\mu U}{r_{0}} \sin \theta$
where $\mathrm{p}_{\infty}$ is the pressure in the absence of motion (or very far from the sphere). The stress vector component on the sphere in the direction of the motion is:

$$
\begin{equation*}
\tau=\tau_{r} \cos \theta+\tau_{\theta} \sin \theta=-p_{\infty} \cos \theta+\frac{3}{2} \frac{\mu U}{r_{0}} \tag{7}
\end{equation*}
$$

from which the net drag on the sphere is found to be:
$D=\int_{0}^{2 \pi} \int_{0}^{\pi} \pi r_{0}^{2} \sin \theta d \theta d \phi=6 \pi \mu U r_{0}$,
which was used to derive the Stokes settling speed. The next order correction for finite Re number gives a modified drag coefficient: $D=6 \pi \mu U r_{0}(1+3 R e / 8)$.

## References:

Acheson, D. J., 1990. Elementery fluid dynamics. Oxford Press.
Batchelor, G. K., 1967. An introduction to fluid dynamics. Cambridge University Press.
Karp-Boss, L., E. Boss, and P. A. Jumars, 1996. Nutrient fluxes to planktonic osmotrophs in the presence of fluid motion. Oceanography and Marine Biology: An annual Review, 34, 71-107.

