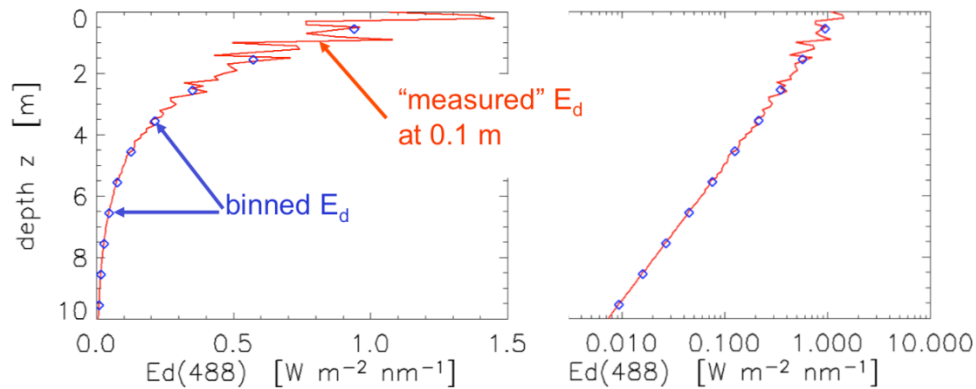


Cruise data analysis: Radiometric properties and AOPs.

Processing assignments:

Generate dark corrected and surface irradiance reference corrected (profiled measurements only) radiometric quantities. Use the median surface irradiance over the cast for the latter normalization. Filter out high tilts (e.g., > 5 deg) and all upcasts in the profiled data. Compute surface PAR and PAR(z) from these radiometric quantities. Generate several depth-binned versions of the profiles using a different vertical bin size for each (from 20 cm to 2 m). Visually verify that the vertically binned data accurately represent the un-binned profiles. Subsample the depth profiles further to simulate data collected on a buoy or mooring by retaining data at only 1, 5, and 9 m.



Assignments:

For simplicity and clarity, you may choose to focus on three RGB wavelengths, e.g., 443, 555, and 670 nm. Derive $K_d(z)$, $K_{Lu}(z)$, and $K_{par}(z)$ for each of the vertically binned profiles, including the 1/4/9 m mooring simulation. How do they relate? Are $K_d(z)$ and $K_{Lu}(z)$ similar? Are $K_d(z)$ and $K_{par}(z)$ similar? Using a single method to calculate K , derive average K_d , K_{Lu} , and K_{par} for three water column thickness: the very near surface layer, the first optical depth (37% light level), and the second optical depth (13% light level). How do they relate? For one of the vertical profiles, re-derive these average values using different calculation techniques: averaging $K_d(z)$, linear fitting to $\log(Ed)$ vs. z , and a nonlinear fitting to Ed vs. z . How do they relate? For the latter two, recall that $Ed(z) = Ed(0^-) \exp(-K_d z)$ and that you did very similar curve fitting in your absorption and attenuation labs.

Calculate L_w , $Ed(0^+)$, and R_{rs} for the HyperPRO profiles. Do so using vertically binned and un-binned data. How do the R_{rs} compare? Over which depth intervals did you calculate K 's to extrapolate to near-surface values? Do these values have spectral dependence? Do you get similar R_{rs} if you use K 's calculated over larger depth intervals (e.g., first and second optical depths)? (Unfortunately, recall that the HyperPRO data below 22 m cannot be used because of a problem with the pressure sensor.) STILL

RELEVANT? How do the $E_d(0+)$ extrapolated from $E_d(z)$ compare with $E_d(0+)$ from the surface reference sensor?

Generate a profile where you haven't filtered for high tilt. How do the derived K 's and R_{rs} compare? Calculate R_{rs} using profiles collected in the ship shadow and with an inaccurate pressure tare. How do the derived K 's and R_{rs} compare? Generate a profile (one with a patchy sky, probably from Cruise 1) where you haven't normalized the in-water data to the reference surface irradiance. How do the derived K 's and R_{rs} compare? For all of the above, how do these $E_d(0+)$ extrapolated from $E_d(z)$ compare with $E_d(0+)$ from the surface (above-water) reference sensor? Comment on the need for accurate pressure tares and knowledge of the offsets between the L_u , E_d , and pressure sensors.

Calculate L_w and R_{rs} for the HyperSAS, Immersed L_u , and Lee method. How do these values compare with those from the vertical profile? Did either the HTSRB or HyperSAS observe changes in R_{rs} with time? Do you see tilt and roll effects in the Immersed data? Can you identify times when the HyperSAS was pointed too close to the plane of the sun (e.g., < 90 deg azimuth)?

Calculate $a(z)$, $b(z)$, $c(z)$, and $bb(z)$ for the slow-drop and wired profiles (note: add salt water contribution). How do they compare? K_d is proportional to $(a + bb)$. Model $K_d(z)$ using your IOPs – for example, following Gordon 1989, which describes $K_d = 1.0395 * (a + bb) / \mu_w$, where μ_w is $\cos(\text{solar zenith angle in water})$. Other methods include Kirk 1984, Morel and Loisel 1998, and Lee et al. 2005. How does $K_d(z)$ derived from AOPs compare with that derived from IOPs? Calculate optically-weighted, average IOPs for the first optical depth using the appropriate K_d (following, e.g., Smith 1981). How do IOPs collected at the surface and near a deep Chl maximum compare with depth-weighted IOPs? How do the median IOPs for the first optical depth compare with their optically weighted averages?